



Mid-Latitude Atmospheric Circulation, Subtropical Regimes, and ENSO

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Some issues regarding the theory of the General Atmospheric Circulation

- The stability properties of the time mean state do not approximate the dynamical properties of the full nonlinear system
- Impossibility of creating a self-consistent theory of the time-mean circulation relying only on the time-mean fields.
- Impossibility of using straightforwardly the fluctuation-dissipation theorem, since ergodicity is only restricted to the attractor, and the identification of Atlantic blocking, to recent work on regimes detection and identification.
- Externally induced fluctuations move the system out of the attractor with probability 1 (Gallavotti/Cohen)
- Impossibility of parameterizing exactly a Climate Change theory: external and internal fluctuations are not equivalent.
- In the limit of infinite resolution for any numerical model of fluid flow, the numerical convergence to the statistical properties of the continuum real fluid dynamics is not guaranteed
- Models of ever-increasing resolution may not give the ultimate answer.

Goals of this Work

- The notion that well-defined winter mid-latitude atmospheric patterns of flow are recurrent during the NH winters has been debated since the definition of Grosswetterlage, to the identification of Atlantic blocking, to recent work on regimes detection and identification.
- Planetary waves contribute to a large portion of the eddy meridional heat transport \Rightarrow good representation by GCMs is needed.
- A widely accepted theory for planetary waves is not available (bimodality?) and reliable data are needed to test new theories.
- Understanding the properties of planetary waves may help in addressing the feasibility of extended range forecasts or the robust detection of climate changes.
- This work is part of the IPCC sub-diagnoses project *Regimes of the Mid-latitude Atmosphere* by Speranza and Lucarini.

Theoretical Background

- At first, the zonal flow – wave field interaction (via form-drag) was proposed as driving mechanisms allowing for the establishment of multiple equilibria of the planetary waves amplitude (Charney and DeVore, 1979)
- However, the transitions between the quasi-stable equilibria require for energetic reasons large variations ($\Delta U \approx 40 \text{ m/s}$) of the mean westerlies, at odds with the “normality” of the distribution of the observed westerlies strength
- The bent resonance curve obtained in non-linear wave self interaction theories by Benzi et al. [1986], not only explains the existence of the multiple equilibria of the planetary wave amplitude, but allows that relatively small changes of the jet strength may imply a switch from unimodal to multimodal regimes of the atmospheric circulation \Rightarrow Baroclinic processes are relevant for the onset and maintenance of planetary waves, and that topography might be crucial.

Data and Methods

Two proper counterparts to the dynamical parameter employed in the theories are extracted from such data sets by computing two robust indicators of relevant large scale features of the midlatitude troposphere. The Wave Activity Index (WAI), introduced by Hansen and Sutera [1986], is computed as the root mean square of the zonal wavenumbers 2 to 4 variance of the winter 500 hPa geopotential height averaged over the channel 32°N - 72°N . The WAI provides a synthetic picture of the ultra long planetary waves and captures the orographic resonance, since an approximate mode of zero phase velocity (resonance) is 3. The Jet Strength Index (JSI) is computed daily as the maximum of the zonal mean of the zonal wind at 200 hPa, where the subtropical jet peaks.

$$WAI = \left(\sum_{k=2}^4 2|Z_k|^2 \right)^{\frac{1}{2}} \quad JSI = \max \langle U \rangle \Big|_{200 \text{ hPa}}^{32^\circ\text{N}-72^\circ\text{N}}$$

For filtering out the synoptic atmospheric variability, we apply a low-pass filter to both WAI and JSI indexes, whereas, for capturing the anomalies with respect to the seasonal cycle we filter out from the WAI signal the spectral peaks occurring at 12, 6 and 4 months, directly related to influence of the external solar forcing. Since in the context of the jet resonance theory the jet strength is considered as an autonomous forcing parameter of the system, controlling and catalysing its internal variability, we do not filter out from the JSI the seasonal cycle and its harmonics as done for the WAI. Furthermore, the results are robust with respect to the different filtering techniques.

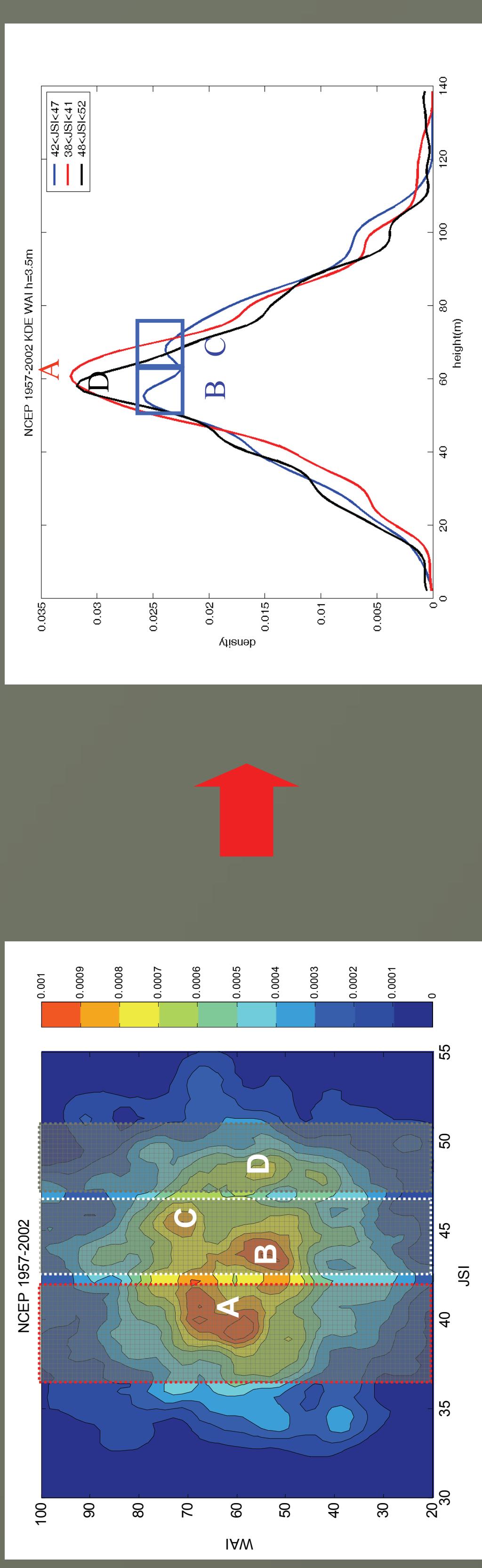
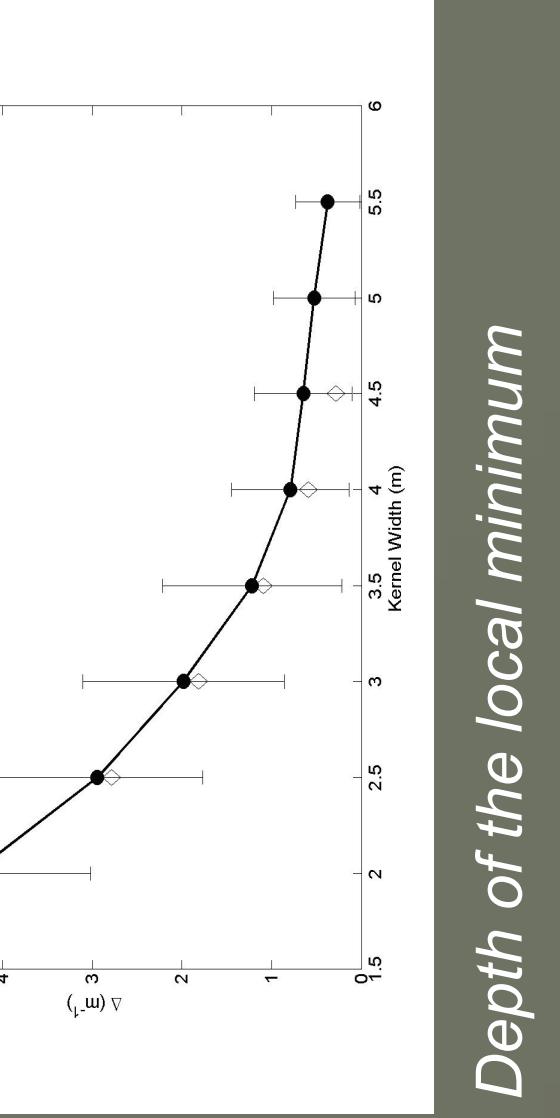
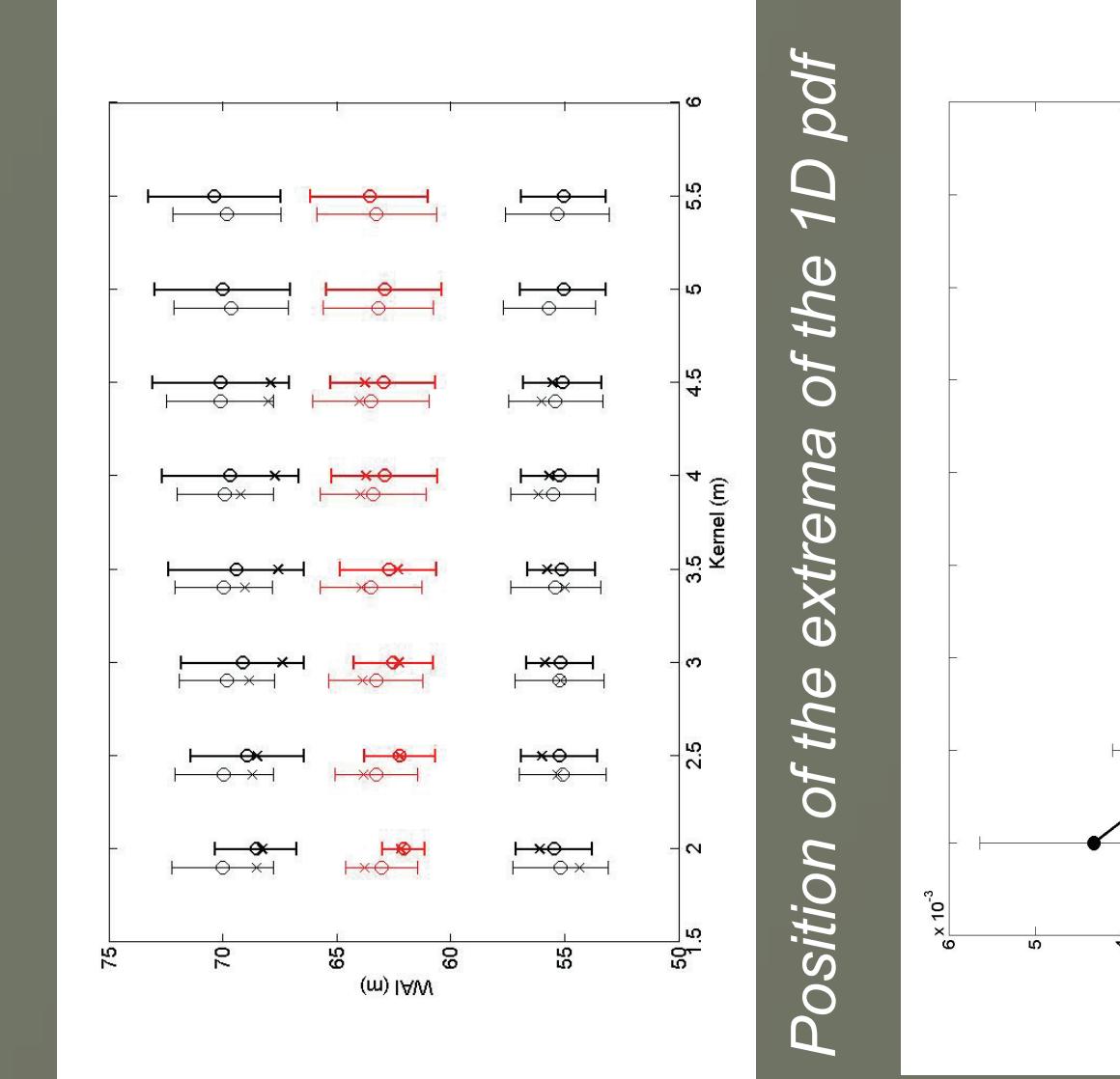
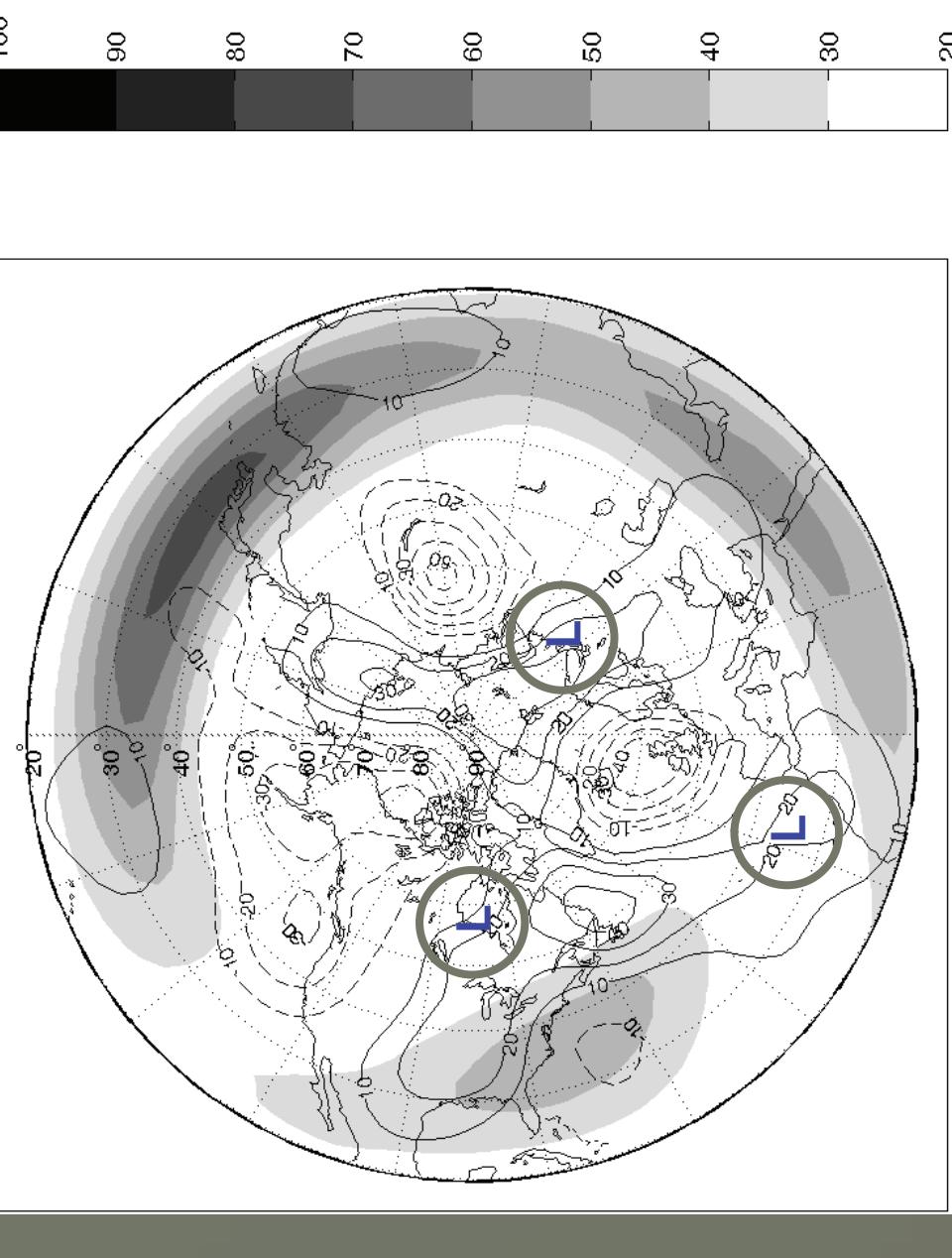
Reanalyses:

- We consider the 1958-2002 DJF [JSI,WAI] time series derived from the NCEP/NCAR and ERA40 reanalysis daily data. Both the 2D [WAI,JSI] and the 1D WAI and JSI time series of the two datasets are determined as equivalent by the 2D and 1D Kolmogorov tests, respectively.
- IPCC 4AR GCMs:

- We also consider the 1961-2000 simulations of the MIROC(hires) and GFDL2.1 models, which have been shown to provide the best representation of NH mid-latitude winter atmospheric variability. When considering GCMs data, we face the difficulty that Z levels are not included in the PCMDI dataset!
- Option 1: Z500 reconstruction via hydrostatic averaging (we need only the anomalies!), using the geostrophic relationship.
- Option 2: Geostrophic relation at 500hPa + Latitudinal averaging (we need the JSI and WAI values).
- We choose Option 2 and obtain the 1961-2000 DJF [JSI,WAI] time series from the two GCMs. The time-series generated by the two models are not equivalent between themselves and are not equivalent to those of the reanalyses.

Results

The empirical 2D joint pdfs, constructed by means of 2D Gaussian estimators, present multiple, well defined, peaks. NCEP results are shown. A major peak (A) corresponds to weak upper tropospheric jet ($JSI = 40 \text{ ms}^{-1}$) and low-to-intermediate activity of the planetary waves ($WAI = 60 \text{ m}$). A second peak (B) corresponds to intermediate values of JSI and low values of WAI ($JSI = 43 \text{ ms}^{-1}$; $WAI = 55 \text{ m}$). The third peak (C) corresponds to similar intensities of the jet ($JSI = 45 \text{ ms}^{-1}$) and very high values of ($WAI = 70 \text{ m}$). The fourth peak (D), corresponding to intense jet ($JSI = 50 \text{ ms}^{-1}$), features relatively weak planetary waves ($WAI = 55 \text{ m}$).



NCEP 1957-2002 DJF WAI=60.5m

gfdl2.1 JSI=44.5

45JSI=48

40.5JSI=43

40JSI=47

35JSI=45

35JSI=48

30JSI=46

30JSI=47

25JSI=48

25JSI=49

20JSI=50

20JSI=51

15JSI=52

15JSI=53

10JSI=54

10JSI=55

5JSI=56

5JSI=57

0JSI=58

0JSI=59

5JSI=60

10JSI=61

15JSI=62

20JSI=63

25JSI=64

30JSI=65

35JSI=66

40JSI=67

45JSI=68

50JSI=69

55JSI=70

60JSI=71

65JSI=72

70JSI=73

75JSI=74

80JSI=75

85JSI=76

90JSI=77

95JSI=78

100JSI=79

105JSI=80

110JSI=81

115JSI=82

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